The Extragalactic Background Light gamma-gamma absorption of Very-High-Energy gamma-rays from cosmological gamma-rays sources

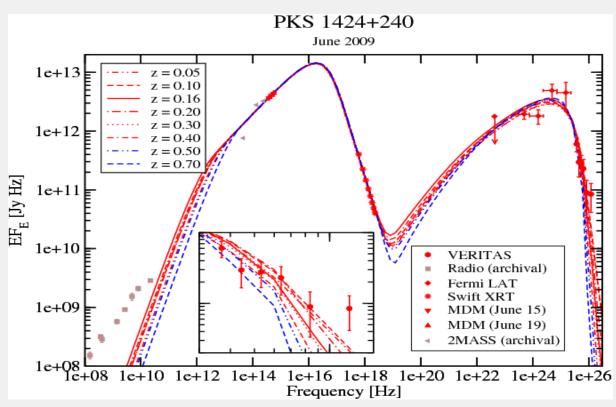
Hassan Abdalla Supervisor: Prof. M. Böttcher



Very High Energy Gamma Rays (VHE; more than 100 GeV) from Cosmological Gamma Ray Sources such as Blazars can be absorbed by the Extragalactic Background light (EBL).

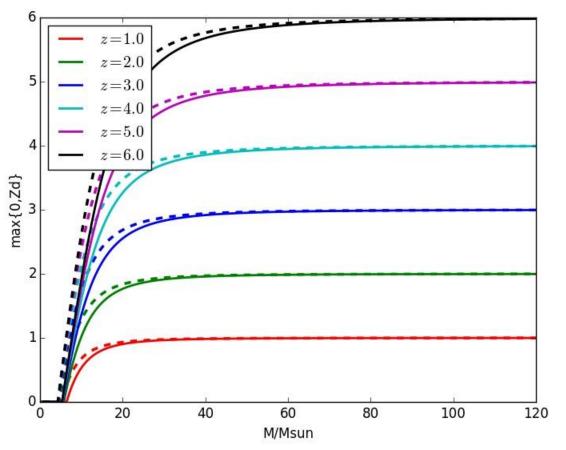
This leads to a high-energy "cut-off" at the VHE end of Blazar spectra.

- The probability of absorption depends on the photon energy and redshift.
- This process has been intensively studied during the last few decades (e.g., Stecker 1969 .... Aharonian et al. 2006).



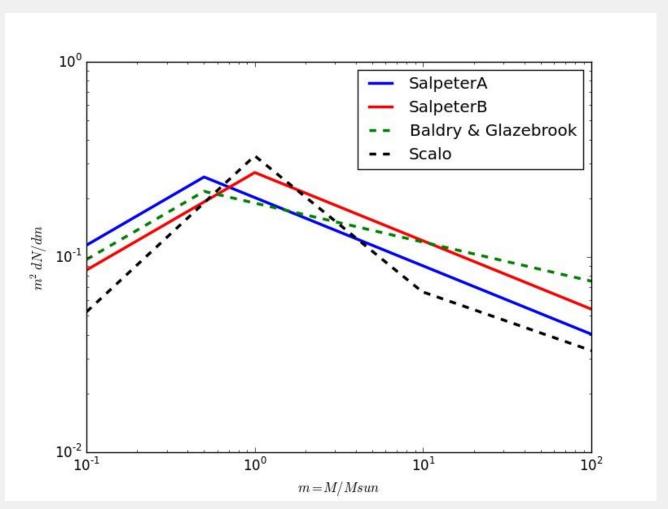
Star from its birth at redshift z to the redshift which it has evolved away from the main sequence  $Z_d(M,z)$ 

$$Z_d(M,z) = -1 + \left(-\frac{\Omega_{\Lambda}}{\Omega_m} \operatorname{sech}\left[\frac{3}{2}H_0t_* + \tanh^{-1}\sqrt{1 + \frac{\Omega_m}{\Omega_{\Lambda}}(1+z)^3}\right]^2\right)^{\frac{1}{3}}.$$



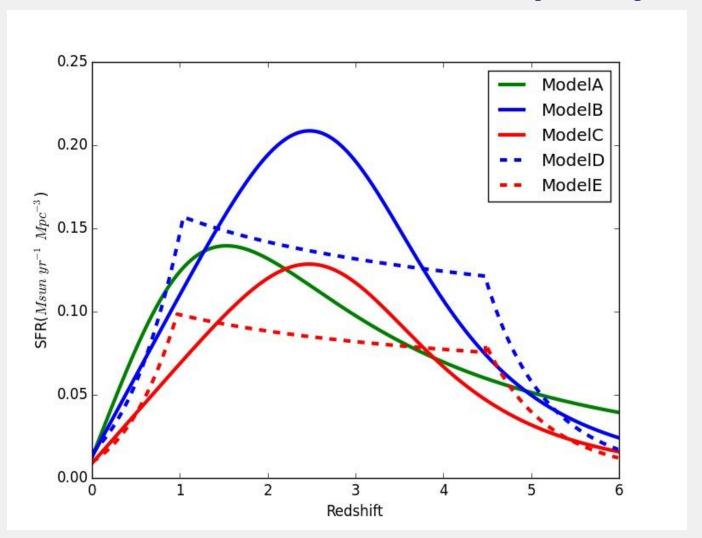
Reproduction from Razzaque et al. (2009)

#### *Initial Mass Function (IMF):*



Reproduction from Razzaque et al. (2009)

## Star Formation Rate (SFR):



The integrated number of photons that have been emitted from star birth to the present epoch  $\frac{dN(\epsilon,M)}{d\epsilon}$ 

$$\frac{dN(\epsilon,M)}{d\epsilon} = \int_{\max\{0,Z_d(M,z')\}}^{z} dz' \frac{dt}{dz'} \frac{dN(\epsilon',M)}{d\epsilon' dt} (1+z')$$

$$\frac{dN(\epsilon',M)}{d\epsilon' dt} = \pi R^2 c \frac{dN}{d\epsilon' dV}$$

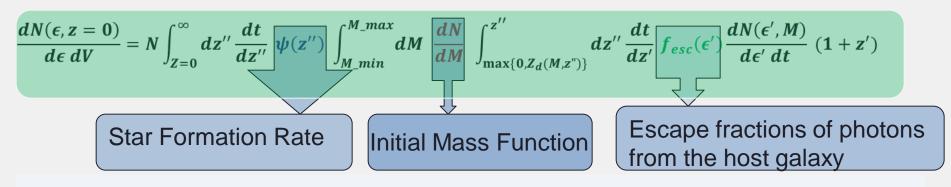
$$\frac{1}{-H_0(1+z)\sqrt{\Omega_m(1+z)^3 + \Omega_\Lambda}}$$
Total number of photons

thermal Black Body photons

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emitted per unit energy

and time interval



Note: The EBL energy density measured at redshift z = z1 can be transformed from z = 0by calculating the comoving energy density

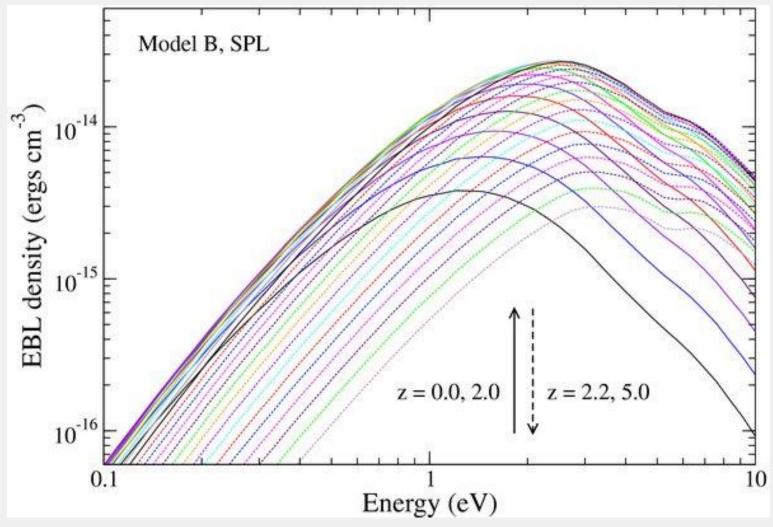
and volume as: 
$$\epsilon_1 \Rightarrow \epsilon(1+z1)$$
 and  $V_1 \Rightarrow \frac{V}{(1+z1)^3}$ 

$$\epsilon \mu_{\epsilon} = (1+z)^4 \epsilon^2 \frac{dN(\epsilon,z)}{d\epsilon dV}$$

$$\epsilon \mu_{\epsilon} = (1+z)^4 \epsilon^2 \frac{dN(\epsilon,z)}{d\epsilon dV}$$

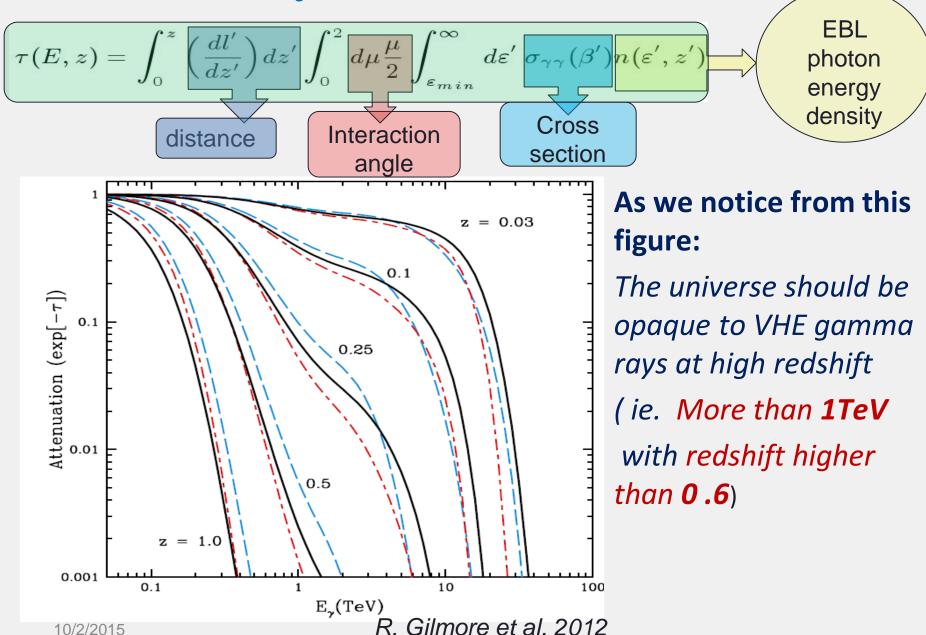
Razzaque et al. (2009)

# The comoving photon energy with different redshift z for EBL model



## Gamma-Ray attenuation:

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# **But from recent Observation:**

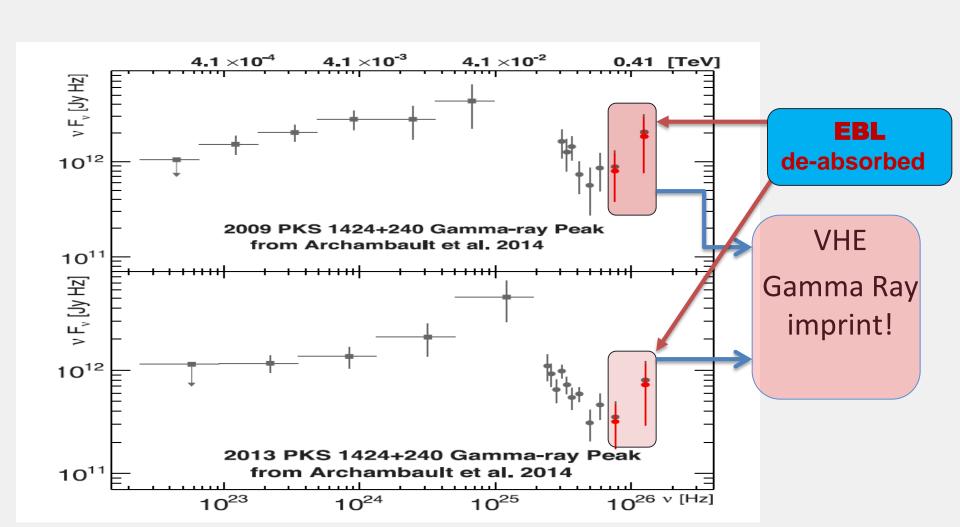
# The universe is more transparent to the VHE-Gamma Rays

than was expected!

Archambault et al. 2014

## Inhomogeneities in the EBL distribution:

Furnis et al., found possible correlation between detected VHE gamma rays sources and cosmic void along the line of sight!



# TO EXPLAIN THIS VHE SIGNATURES IN THE SPECTRA OF DISTANT BLAZARS

### There are Possible solutions:

Existence of exotic Axion Like Particles (ALPs)

Dominguez et al. 2011

 Interactions of extragalactic Ultrahigh Energy Cosmic Rays (UHECR)

Essey et al. 2010

 The existence of cosmic voids between the Blazar and the observer on the earth, which means the EBL not homogeneous.

Furniss et al. 2013

# Inhomogeneity in the EBL distribution:



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# Summary:

# This problem also maybe due to:

- Misunderstanding of the emission mechanisms in the Blazars and/or the propagation of VHE photons through the Inter Galactic Meduim.
- EBL physics not well understood

In our current work we are trying to study the expected inhomogeneity of the EBL, in particular if the line of sight to a Blazar is passing through large voids in intergalactic space!

# Thank you !!!

